# LETTERS TO THE EDITOR

## CONVERSION OF LII AND LIII ELECTRONS WITH EMISSION OF E1-QUANTUM IN E0 NUCLEAR TRANSITIONS

### D. P. GRECHUKHIN

Academy of Sciences, U.S.S.R.

Submitted to JETP editor May 5, 1957

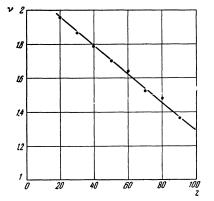
J. Exptl. Theoret. Phys. (U.S.S.R.) 33, 1037-1038 (October, 1957)

 $S_{\text{HELL}}$  electron conversion in nuclear E0 transition is determined by the interaction in the region  $r \leq R_0$  inside the nucleus. The probability of E0 conversion of the LII and LIII electrons is therefore substantially less than that of the LI electron (see Ref. 1).

$$W_{E_0}(L \text{ II}) = \frac{3}{2} (Ze^2/2)^2 p_f^2 (\varepsilon_f + 1)^{-2} W_{E_0}(L \text{ I});$$
  
$$W_{E_0}(L \text{ III}) \sim (p_f R_0)^2 (Ze^2 R_0/2)^2 W_{E_0}(L \text{ I}) \approx 10^{-8} \div 10^{-9} W_{E_0}(L \text{ I}).$$

Here  $p_f$ ,  $\epsilon_f$  are the electron energy and momentum in the final state; relativistic units  $\hbar = m = c = 1$ and  $e^2 = 1/137$  are used throughout.

It would be interesting to consider conversion of LII and LIII electrons whereby the electron, emitting a dipole quantum (E1), changes to a state with momentum  $j = \frac{1}{2}$ ,  $\ell = 0$  and then converts in the E0-



nuclear transition. We give below the results of the calculations. We consider the emission of quanta with energies  $k \leq Ze^2/2$ ; the calculations are carried out with the Coulomb functions of the electrons in the field of the nucleus in second-order perturbation theory. Screening is neglected. The finite dimensions of the nucleus are taken into account only in the calculation of the matrix element of the E0 transition.

After inserting the matrix elements of the E0 and E1 transitions, the integrals over the intermediate states for LII and LIII electrons reduce to integrals of the type

$$J = \int_{0}^{\infty} dx \frac{\Phi_{\delta}(x) x^{m}}{(1+x^{2})^{n} (1+x^{2}+2\xi/\alpha)}, \quad x = \frac{p}{\alpha}, \quad m < n.$$

where

$$\mathbf{x} = Ze^2/2, \ \Phi_{\delta}(\mathbf{x}) = x^{2\gamma-1} \left(1 + x^2\right)^{3\delta/2} \exp\left(-4\varepsilon \ \tan^{-1} x/x\right), \quad \gamma = \sqrt{1-4\delta}, \quad \delta = \alpha^2, \quad \xi = k/\alpha$$

To calculate this integral,  $\Phi_{\delta}$  is approximated in the interval  $0.2 \le x \le 15$  by means of a power function  $e^{-4+\lambda}x^{\nu}$ . The dependence of  $\nu$  on Z is given in the diagram. The value of  $\lambda$  is restricted to  $0.9 \le \lambda \le 1.1$ . Thus the permissible error does not exceed 20%. The calculated value of J is

$$J = \frac{\exp\left(-4+\lambda\right)}{2} I_n^{\nu m} \left(-\frac{2\xi}{\alpha}\right) = \frac{\exp\left(-4+\lambda\right)}{2\Gamma\left(n+1\right)} \Gamma\left(n+\frac{1-m-\nu}{2}\right) \Gamma\left(\frac{1+\nu+m}{2}\right) {}_2F_1\left(1, n+\frac{1-m-\nu}{2}; n+1; -\frac{2\xi}{\alpha}\right),$$

from which we obtain for the unknown probabilities

$$dW_{\gamma E_0} (L \text{ II}) = 1.2 \cdot 2^{4\gamma - 2} \left[ \Gamma \left( 2\gamma + 1 \right) \right]^{-2} e^2 W_{E_0} (L \text{ I}) \left\{ \left[ I_2^{\nu_0} + \frac{\gamma + 1}{6} \xi^2 \left( \frac{2\gamma + 3}{3} \mathfrak{M} + 2I_3^{\nu_0} - (2\gamma + 3) \mathfrak{R} \right) \right]^2 \right\} \\ + 2 \left[ \frac{(2\gamma + 3)(2\gamma + 2)}{45} \right]^2 \xi^4 \mathfrak{M}^2 \left\} \delta \left( \Delta - \varepsilon_1 - k - \varepsilon_f + 1 \right) \xi d\xi d\varepsilon_f, \quad \mathfrak{M} \approx I_5^{\nu_0} - I_5^{\nu_4} - 4I_5^{\nu_2}, \quad \mathfrak{M} \approx 0.6I_7^{\nu_5} + 2.5I_7^{\nu_6} - 5I_7^{\nu_2} - 3I_7^{\nu_0};$$

$$dW_{\gamma E0} (L \text{ III}) = 2.2 \frac{2^{4\gamma - 2} [\Gamma (\gamma_1 + \gamma + 1)]^2}{\Gamma (2\gamma_1 + 1) [\Gamma (2\gamma + 2)]^3} W_{E0} (L \text{ I}) \left\{ \left[ I_2^{\nu_0} + \frac{(\gamma + 3)(\gamma + 4)}{6} \xi^2 \mathfrak{M} \right]^2 \right\}$$

$$+\frac{9}{32}\left[\frac{(2\gamma+1)(\gamma+3)(\gamma+4)}{15}\right]^{2}\xi^{4}\left(\Re+\frac{5}{3}\Re\right)^{2}\right]\delta\left(\Delta-\varepsilon_{1}-k-\varepsilon_{f}+1\right)\xi d\xi d\varepsilon_{f}, \quad \gamma=\sqrt{1-\alpha^{2}}, \quad \gamma=\sqrt{1-4\alpha^{2}},$$

where  $\Delta$  is the nuclear transition energy and  $\epsilon_1$  is the binding energy of the L electron, equal to  $(\text{Ze}^2)^2/8$ .

If  $\Delta \gg \alpha^2$  it is easy to estimate the lower limit of the total transition probability:

$$W_{\gamma E_0}(L \text{ II}) \gg W_{E_0}(L \text{ I}, \Delta) \frac{1}{137} 1.2 \frac{2^{4\gamma-2}}{[\Gamma(2\gamma+1)]^2} \frac{\alpha^2}{16} C, \quad W_{\gamma E_0}(L \text{ III}) \gg W_{E_0}(L \text{ I}, \Delta) \frac{1}{137} 2.2 \frac{2^{4\gamma-2}[\Gamma(\gamma_1+\gamma+1)]^2}{[\Gamma(2\gamma_1+1)[\Gamma(2\gamma+2)]^3} \frac{\alpha^2}{16} C,$$

where the constant C is  $\gtrsim 2$ .

For the ratio of the two probabilities we obtain

 $W_{E_0}(L \text{ II}) / W_{\gamma E_0}(L \text{ II}) \leq (1.37 \cdot 10^3 / C) 2^{3-4\gamma} [\Gamma (2\gamma + 1)]^2 p_f^2 (\varepsilon_f + 1)^{-2}.$ 

if  $\Delta \sim \alpha^2$  the ratio of the probabilities can be on the order of unity, owing to the smallness of  $p_f$ . The ratio  $W_{E_0}(LIII)/W_{\gamma E_0}(LIII)$  does not exceed  $10^{-4}$  to  $10^{-5}$  for all values of  $\Delta$ .

Also possible for the LIII electron is a  $\gamma - E0$  conversion with a transition first into the state  $j = \frac{1}{2}$ , l = 1 and emission of a M1 and E2 quantum. In this case, however, the probability  $dW_{\gamma E0}$  is proportional to  $W_{E0}$  (LII) and is considerably less than the probability  $dW_{\gamma E0}$  of a process with emission of an E1 quantum.

<sup>1</sup>D. P. Grechukhin, J. Exptl. Theoret. Phys. (U.S.S.R.) 32, 103 (1957), Soviet Phys. JETP 5, 846 (1957).

Translated by J. G. Adashko 201

## SIGN OF THE SPACE CHARGE ON THE AXIS OF A POSITIVE COLUMN IN A LONGI-TUDINAL MAGNETIC FIELD

#### M. V. KONIUKOV

Tula Pedagogical Institute

Submitted to JETP editor May 11, 1957

J. Exptl. Theoret. Phys. (U.S.S.R.) 33, 1039-1040 (October, 1957)

THE magnitude of the space charge in a positive column, together with the magnitude of the transverse electric field, depends for a given current not only on the temperatures and masses of the particles and on the pressure, but also on the superimposed external longitudinal magnetic fields.<sup>1</sup> To calculate the deviation from the quasi-neutrality on the discharge axis one can use a scheme analogous to that proposed by Ecker.<sup>2</sup> However, since direct measurements of the excess space charge is difficult, it is advantageous to change over to transverse gradients.

The ratio of the transverse gradients near the axis in a magnetic field and without the field is given by

$$E_{r}^{H}/E_{r} = Z^{H}(L_{e}^{H} - L_{p}^{H})(D_{e}b_{p} + D_{p}b_{e})/Z(D_{e} - D_{p})(L_{e}^{H}b_{p}^{H} + L_{p}^{H}b_{e}^{H}),$$
(1)

where  $D_e$ ,  $D_p$ ,  $b_e$ , and  $b_p$  are the coefficients of diffusion and mobility of the electrons and ions without a magnetic field (while the index H denotes the presence of a magnetic field), and Z is the ionization coefficient.

Under conditions usually existing in the positive column of a gas discharge, the transverse field decreases in the presence of a longitudinal magnetic field. The character of the decrease depends on the