quency of 58 mc and acceleration time of 10,000 μ -sec. Preliminary investigation shows the azimuthal spread of the electron bunch at the end of the cycle is 100 $\pm 10^{\circ}$.

A detailed description of the method and results of this experiment will be published at a later date.

The author expresses his sincere gratitude to Prof. P. A. Cerenkov for valuable discussions.

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On the Derivation of a Formula for the Energy Spectrum of Liquid He⁴

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Institute for Physical Problems Academy of Sciences, USSR (Submitted to JETP editor June 10, 1956) J. Exptl. Theoret. Phys. (U.S.S.R.) 31, 536-537 (September, 1956)

A S is well known, Feynman¹, using a wave function of a special form, has obtained an expression for the spectrum of the elementary excitations in liquid He⁴. A hydrodynamical derivation of this formula is presented below.

We shall begin with the Hamiltonian for a quantum liquid in the form²

$$\hat{H} = \frac{1}{2} \int m \hat{\mathbf{v}} n \hat{\mathbf{v}} d\tau + \hat{H}^{1}[n], \qquad (1)$$

where *n* is the number of atoms per unit volume and $\hat{H}^{1}[n]$ is the velocity-independent part of the Hamiltonian. We shall assume it to be a function of *n*. We set $n = \pi + \delta n$ and expand *H* in terms of the second order in δn . The first-order term drops out and we obtain

$$\hat{H} = \hat{H}^{1}[\bar{n}] + \frac{\bar{mn}}{2} \int \hat{\mathbf{v}}^{2} d\tau + \int \varphi(\mathbf{r}, \mathbf{r}') \,\delta n \delta n' d\tau d\tau', \quad (2)$$

where φ is the second functional derivative of \hat{H}^1 with respect to *n*. Transforming now to Fourier components

$$\delta n = \sum_{\mathbf{k}} n_{\mathbf{k}} e^{i\mathbf{k}\mathbf{i}}$$

and taking into account the equation of continuity in the form

$$\delta \dot{n} + \overline{n} \operatorname{div} \hat{v} = 0, \qquad (3)$$

as well as the fact that φ depends only upon |r - r'|, we obtain

$$\hat{H} = \hat{H}^{1}[\bar{n}] + \sum_{\mathbf{k}} \left(\frac{m |\bar{n}_{\mathbf{k}}|^{2}}{2\bar{n}k^{2}} + \frac{1}{2} \varphi_{\mathbf{k}} |n_{\mathbf{k}}|^{2} V \right).$$
(4)

This expression has the form of a sum of the Hamiltonians of oscillators having frequencies:

$$\omega^2(\mathbf{k}) = (k^2 \varphi_{\mathbf{k}} \overline{n}/m) V. \tag{5}$$

For the determination of φ_k we note that the average value of the potential energy of an oscillator in the ground state is equal to $\hbar \omega/4$, whence

$$\frac{1}{2}\varphi_{\mathbf{k}}\overline{|n_{\mathbf{k}}|^{2}}V = \frac{1}{4}\hbar\omega(\mathbf{k}).$$
(6)

As is well known, however, $S(k) = |\overline{n}_k|^2/\overline{n}$ is the Fourier component of the correlation function for the atoms of a liquid, which can be determined from diffraction experiments. Substituting φ_k from (6) into (5), we find for the energy of excitation

$$E(\mathbf{k}) = \hbar \omega (\mathbf{k}) = \hbar^2 k^2 / 2mS(\mathbf{k}),$$

which agrees with Feynman's result.

In conclusion, I would like to express my thanks to L. D. Landau for his advice.

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The Effect of a Transverse Magnetic Field on the Thermal Conductivity of Metals

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L ET us consider a metal, in which there is a heat flow $Q = Q_x$ and a magnetic field $H = H_x$. For the calculation of the coefficient of thermal conductivity we use the model of Sommerfeld,¹ according to which the flow depending on the motion of electrons under the action of the temperature gradient is set equal to zero. Accordingly,²

$$I_{x} = -\frac{3ne}{mv^{3}} \int \xi f v d\varepsilon, \qquad Q_{y} = \frac{3n}{2v^{3}} \int \eta f v^{3} d\varepsilon \quad (1)$$

and analogously for I_y and Q_x . Here (-e) is the charge on the electron, m is the mass of the electron, v is the velocity, ξ and n are the components of the velocity along the x and y axes, ϵ is the kinetic energy of the electron. The distribution function is taken to have the form

$$f = f_0 + \xi \chi_x + \eta \chi_y, \tag{2}$$

where f_0 is the Fermi distribution function, and the functions χ_x and χ_y (found with the aid of the kinetic equation, in which the term taking into account collisions, was derived by Lorentz³) equal:

$$\chi_{x} = -l(f_{1} - qf_{2}) / v(1 + q^{2}),$$

$$\chi_{y} = -l(f_{1}q + f_{2}) / v(1 + q^{2}).$$
(3)

Here l is the length of the mean free path of the electron, and the rest of the variables are defined as follows:

$$q = \omega l / v = (eH / mc) l / v;$$

$$f_1 = \partial f_0 / \partial x - eE_x \partial f_0 / \partial \varepsilon;$$

$$f_2 = \partial f_0 / \partial y - eE_y \partial f_0 / \partial \varepsilon;$$
(4)

 E_x and E_y are the components of the electric field resulting from the motion of the electrons under the action of the temperature gradient.

Calculation shows that the dependence of lon v for the present problem is immaterial, because the terms containing the derivative of l with respect to v, are small and do not enter into the expression for the coefficient for thermal conductivity x. Making the usual calculation for the coefficient

of thermal conductivity $\varkappa = -Q_x / \frac{\partial T}{\partial x}$ (in the present problem $l_x = l_y = 0$, $Q_y = 0$) with accuracy to the terms $\sim (kT/\epsilon)^3$ ($\overline{\epsilon}$ is the Fermi level), we obtain

$$\varkappa = \varkappa_0 \left[1 - \frac{4\pi^2}{15} \left(\frac{kT}{mv^2} \right)^2 \frac{q^2 \left(4 + 2q^2 + 3q^4\right)}{\left(1 + q^2\right)^3} \right],$$
(5)

where $\varkappa_0 = \pi^2 n l k^2 T / 3mv$ is the coefficient of thermal conductivity in the absence of a magnetic field.

Approximate calculation shows that formula (5) gives a decrease in the thermal conductivity of less than 0.01% of κ_0 in a field of 10,000 Oersteds.

It can be shown that consideration of the effect is necessary formetals of the type of Bi which have a small number of conduction electrons.

In conclusion I must thank K. B. Tolpygo for a number of suggestions and E. I. Rashba for certain advice in the course of carrying out the work.

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Application of the Theory of Random Processes to Radiation Transfer Phenomena

L. M. BIBERMAN AND B. A. VEKLENKO Moscow Power Institute (Submitted to JETP editor April 17, 1956) J. Exptl. Theoret. Phys. (U.S.S.R.) 31, 341-342

(August, 1956)

I N this note the motion of the photon is treated as a random process under the following very general assumptions: the medium is isotropic; its properties may be functions of time and space; the photon may be scattered, absorbed by an atom and reemitted, or absorbed in a collision of the second kind; the polarization of the radiation and the motion of the atom excited by a photon are not taken into account.

We begin with the function

 $f_{\nu_1}^{\nu_2}(\mathbf{r}_1, \eta_1, \nu_1, t_1; \mathbf{r}_2, \eta_2, \nu_2, t_2) dV_2 d\eta_2 d\nu_2,$